



Stennis Space Center

ASD FieldSpec FR Calibration

Setup and Techniques

Prepared by

Dan Olive

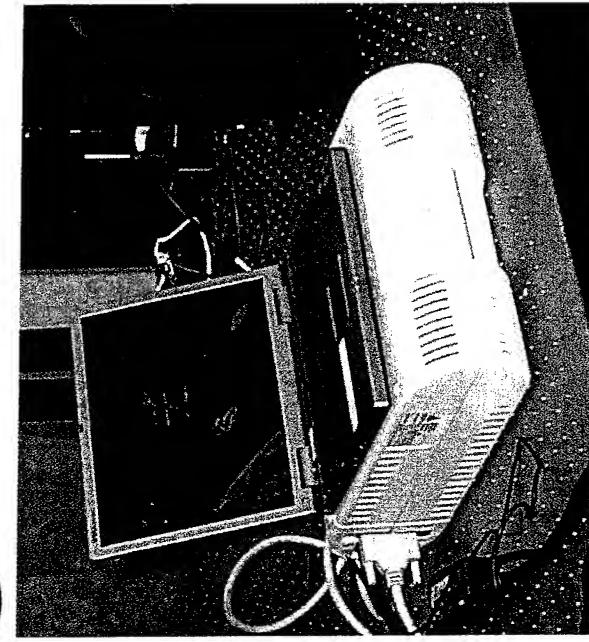
Commercial Remote Sensing Directorate
Lockheed Martin Space Operations – Stennis Programs
John C. Stennis Space Center

October 23, 2001



ASD FieldSpec FR Spectroradiometer

Stennis Space Center



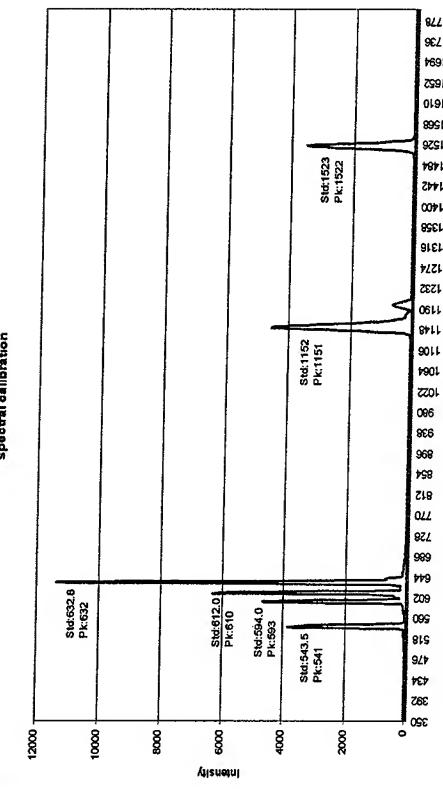
- Designed for collection of spectral data in the field
- Range is 350 to 2500 nanometers
- Measures spectral reflectance, radiance, and irradiance
- Fiber optic bundle carries light to three internal spectrometers
- UV/VIS/NIR silicon array covers about 350 to 950 nm with ~1.4 nm sample interval
- Two NIR/MWIR spectrometers cover 900 to 1850 nm and 1700 to 2500 nm with a sampling interval of about 2 to 3 nm
- Control and data storage with laptop computer
- Interchangeable foreoptics provide flexibility in field of view



Components of Calibration

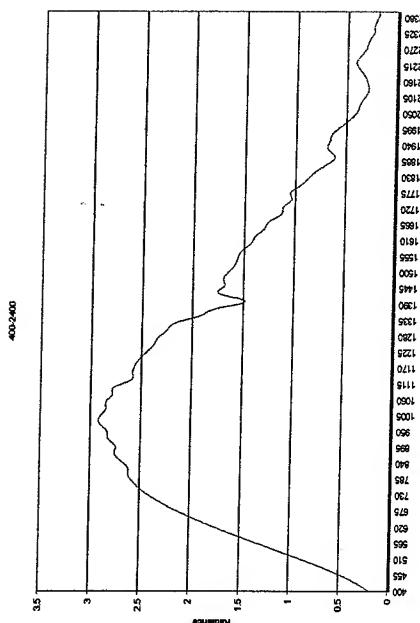
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Spectral calibration



SPECTRAL

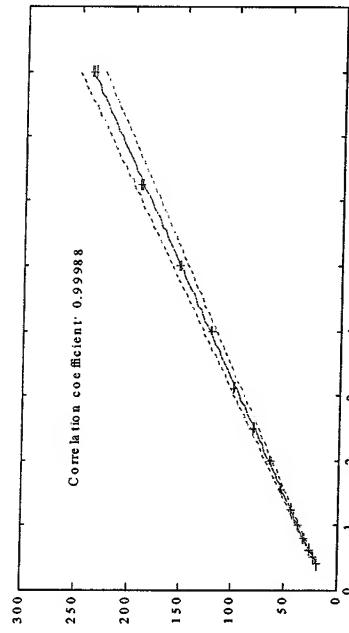
400-2400



RADIOMETRIC

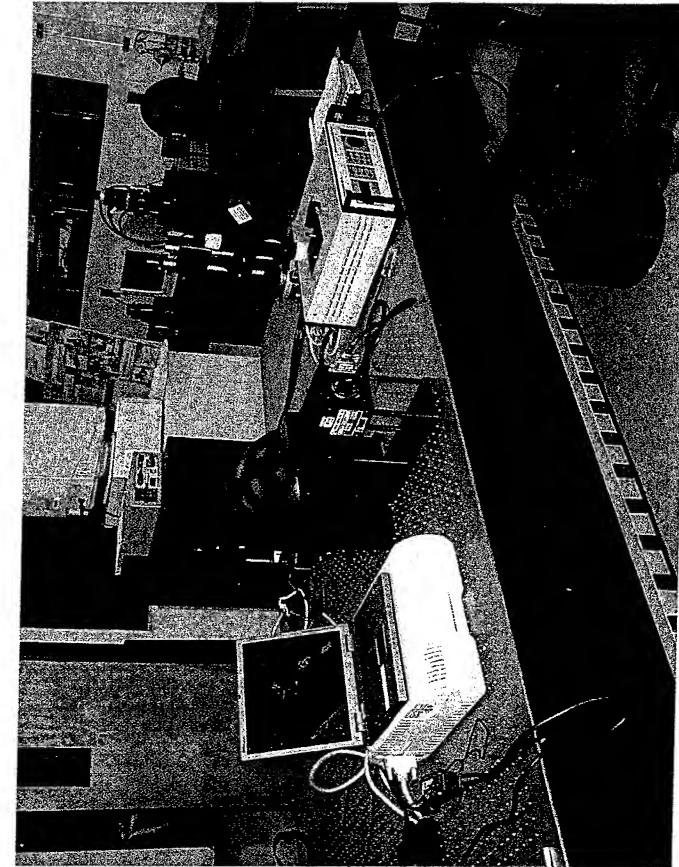
- Spectral: Intrinsic wavelength standards
 - Lasers
 - Discharge lamps

- Radiometric: NIST secondary standard
 - Linearity
 - Field of View (FOV)



LINEARITY

Equipment List

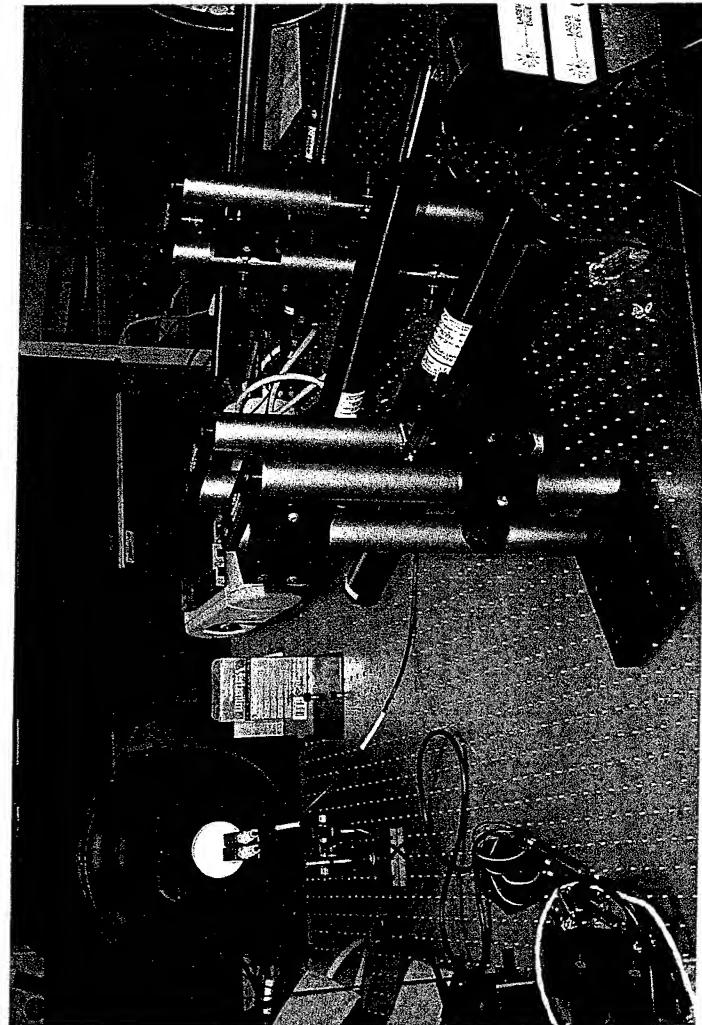


- NIST calibrated integrating sphere
- NIST sphere is equivalent to a secondary radiance standard
- Multiple HeNe lasers
- Sphere controller
- ASD FieldSpec FR spectroradiometer
- Portable laptop computer
- Precision 3-axis positioner
- MATLAB® software



Spectral Setup

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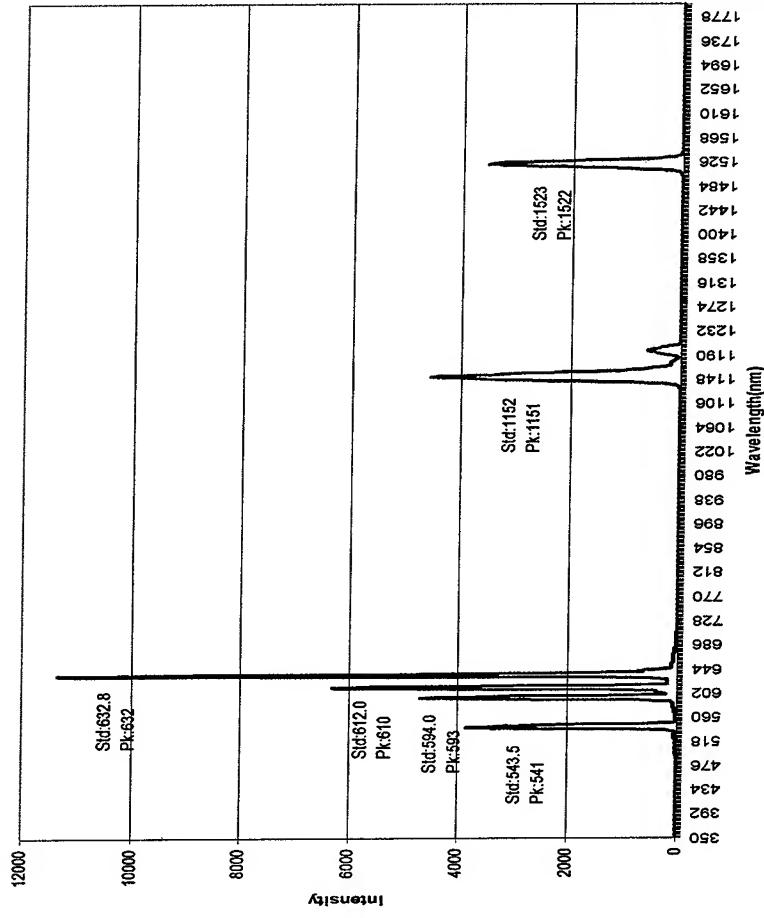
- Use intrinsic spectral sources
- Six different wavelengths
- All HeNe lasers
- Beams directed into sphere
- No beams directly on baffle



Spectral Calibration

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spectral calibration

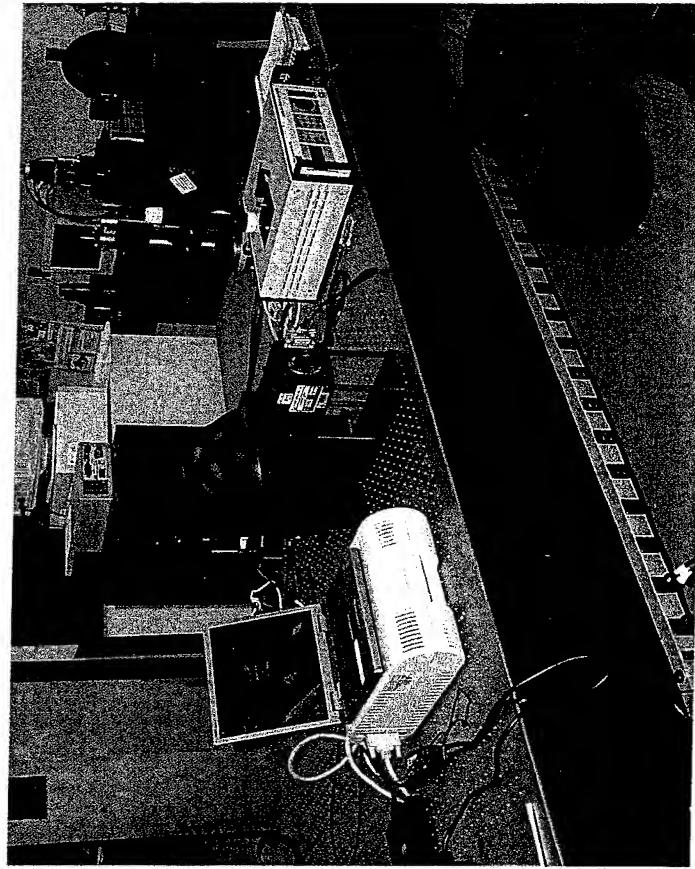


- Take raw DN data with laser input to sphere
- Determine wavelength ASD measures for each intrinsic standard
- Locally generated MATLAB code fits Gaussian to peaks and provides center wavelength and FWHM
- Calculate ASD deviations from accepted standard wavelengths



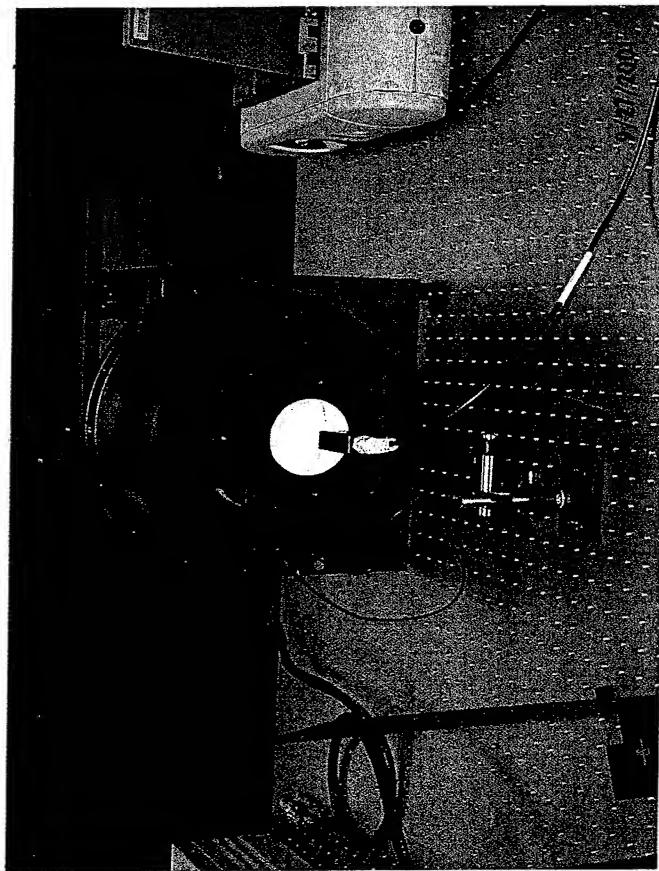
Radiometric and Linearity Setup

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- ASD powered up for 1.5 hours before taking data
- Sphere port set to open away from operator
- Probe with appropriate foreoptic positioned normal to port and centered
- Room lights off for minimum stray light
- No reflective surfaces near output port

Radiometric Setup

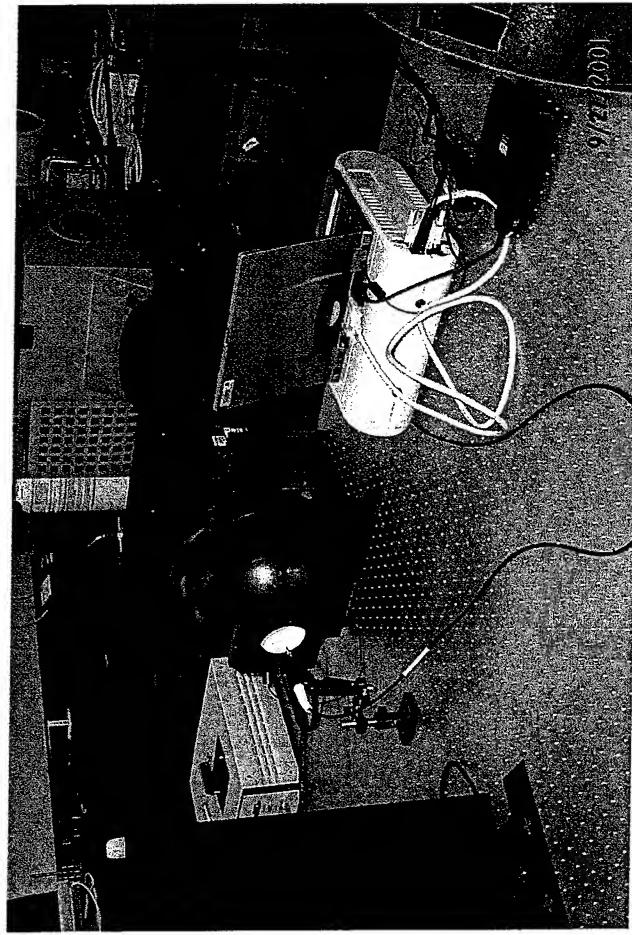


Entrance aperture of probe
close to but not inside port
output plane

- Probe is well centered
- Note edge of internal baffle
- Field of view should be
considerably less than 40° to
ensure total area seen is fully
on the baffle surface



Datasets Required



- NIST calibration at 30.0 mW/cm²*sr
- Equivalent to about 50% of maximum output from sphere
- Linearity requires data collection at 100%, 75%, 50%, 25%, and 10% of max
- Least squares linear fit to these points
- Use data from 50% for radiometric calibration coefficient calculation



Data Files

- All data recorded as raw data files
- Each has the format HEADER.XXXX, where HEADER is created by the operator and XXXX is a number auto-incremented by the FR
- Each file has wavelength scale in one column, DN in a second column
- Files converted to text files (*.txt) with the ViewSpecPro software before analysis
- Calibration Work Instruction details entries so that each experimental data sequence produces a spectrum that is the average of 25 separate readouts
- A total of 30 of these saved spectra is generated for each experimental setup
- First step in analysis is to average these 30 files at each wavelength



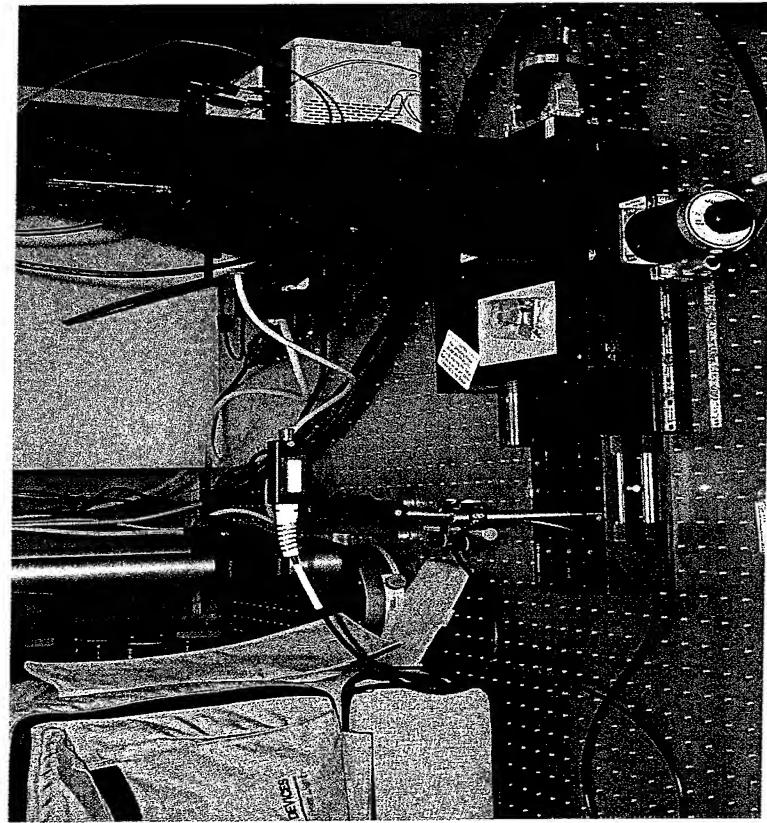
Data Files (continued)

- Averaging 25 readouts for one spectrum reduces noise
- Saving 30 spectra enables statistical calculation of SNR and further reduces noise
- Spectral calibration files of laser lines are analyzed to yield center wavelength and FWHM as measured by the ASD
- Corrections (offsets) are calculated from differences between standard wavelengths and those measured by the ASD
- Spectral radiance calibration coefficients are produced from raw digital numbers recorded and NIST file of spectral radiance from the sphere
- For verification of linearity, the raw spectra are first converted to spectral radiance via the calibration coefficients
- The resultant spectral radiance file is then integrated to yield total integrated radiance, which is used to derive a least squares fit to a line

Field of View Measurement



- Precision, automated three-axis positioner moves small grain of wheat bulb through field of view
- Room darkened
- ASD takes spectrum at each position
- Spectrum is integrated to yield overall intensity and plotted as function of position in the X, Z plane
- Angular FOV reconstructed from geometry





Summary

- **Field portable spectroradiometers play key role in validation and verification**
 - Ground truth data collection
 - Radiometer transfer for internal laboratory use
- **Calibration required under ISO9000**
- **Vendor calibrations not traceable**
- **In-house calibration highly desirable**
 - Traceability ensured
 - Equipment not subject to hazards of shipping
- **Instruments serve multiple purposes**

REPORT DOCUMENTATION PAGE

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